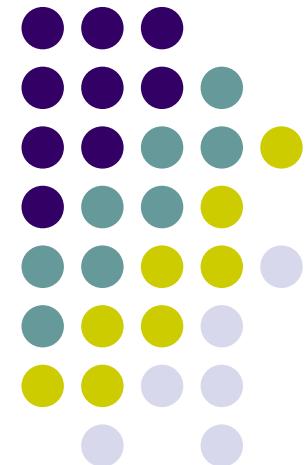


NMR Diffusion Diffraction and Diffusion Interference from Cells

Philip Kuchel & Guilhem Pages



School of Molecular and Microbial Biosciences University of Sydney

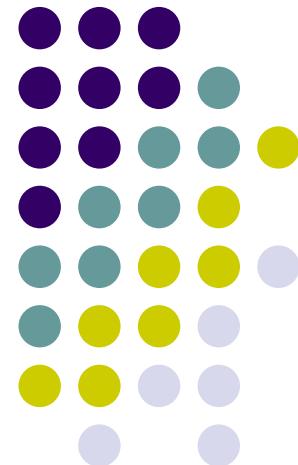


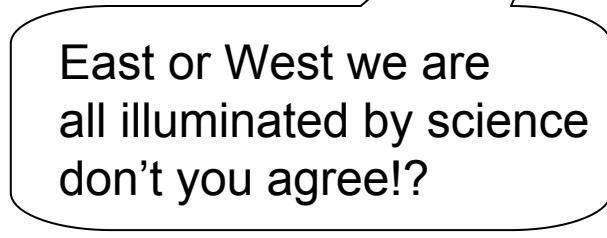
Related posters

D2 Levitz, P. Intermittent Brownian dynamics over strands

E2 Gratz, M. and Galvasos, P. Methodical aspects of 2D NMR correlation spectroscopy under conditions of ultra high pulsed field gradients

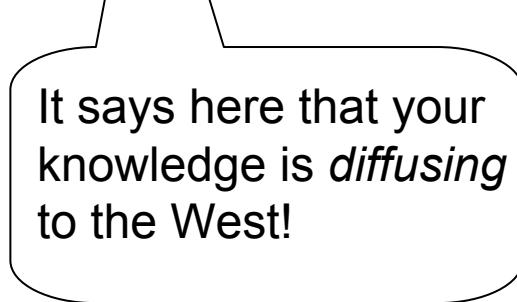
E3 Veil, S., Excoffier, G., Pages, G., Ziarelli, F., Delaurant, C., and Caldarelli, S. Combined use of pulsed gradient spin echo and high resolution magic angle spinning to investigate solute diffusion in the presence of chromatographic stationary phase





East or West we are
all illuminated by science
don't you agree!?

QuickTime™ and a
TIFF (Uncompressed) decompressor
are needed to see this picture.



It says here that your
knowledge is *diffusing*
to the West!



Contents

- Red cells...motivation
- q -Space analysis
- Flow diffraction
- Octagon-star model



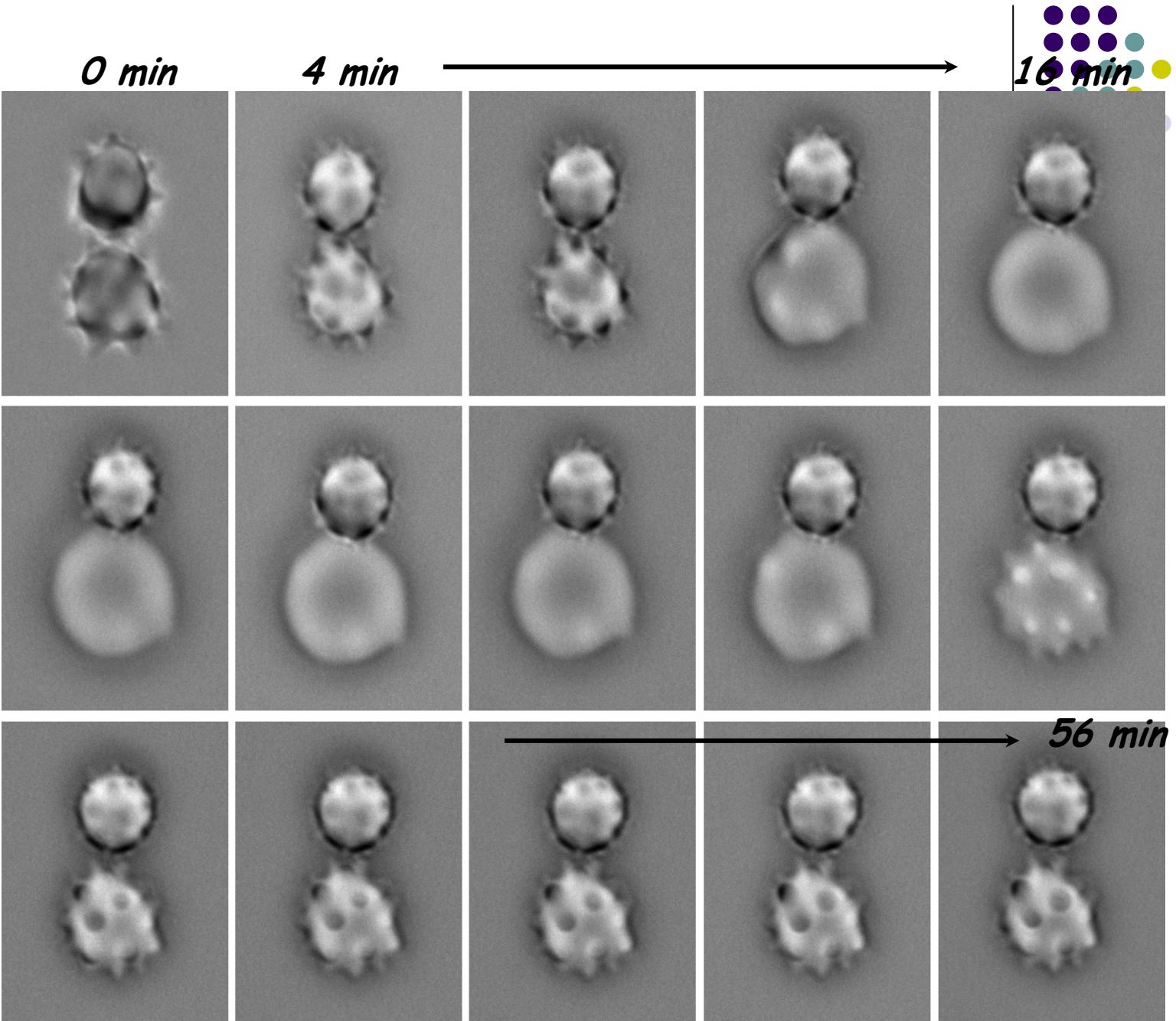
QuickTime™ and a
DivX 4.1.2 decompressor
are needed to see this picture.

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*Echinocytes
under DIC
Microscopy*

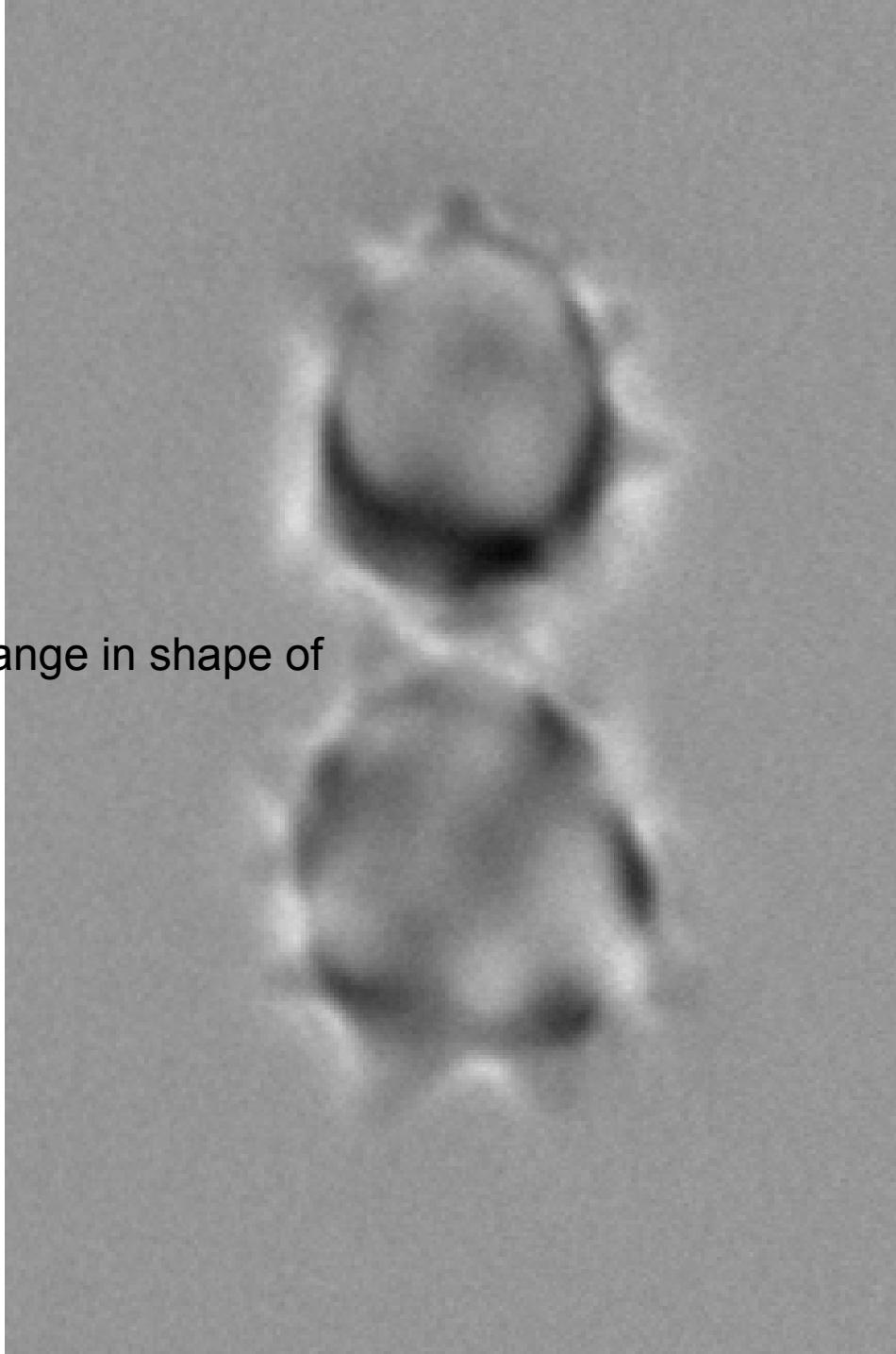
*...showing
echinocyte
to discocyte
reversion
and back
again*





*Echinocytes
under DIC
microscopy*

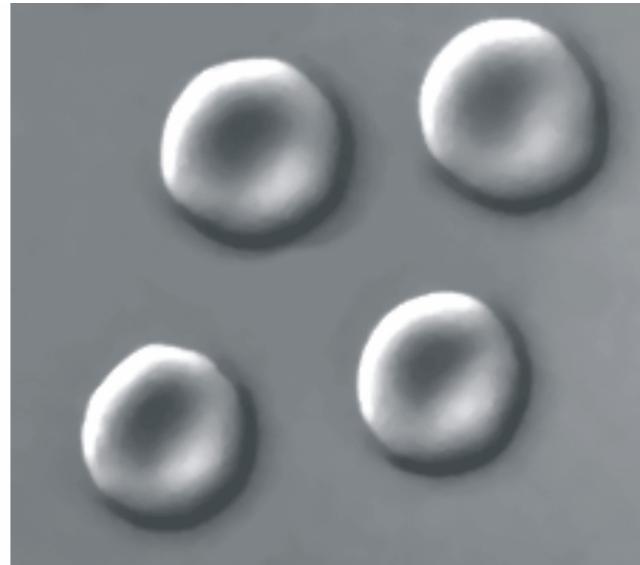
Movie showing change in shape of
red blood cell





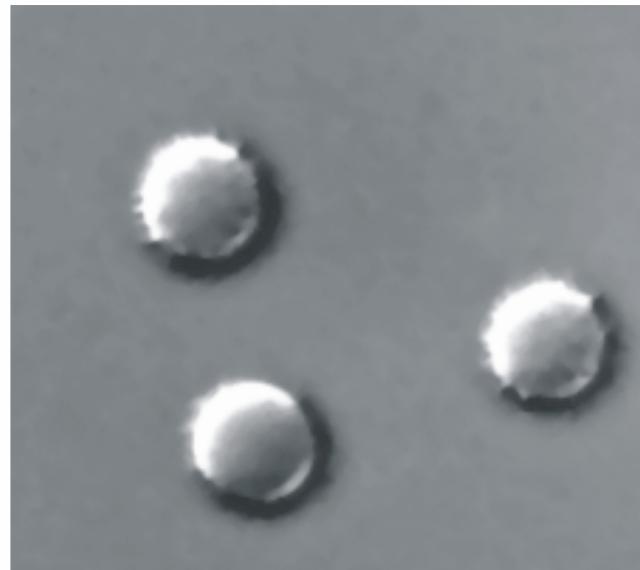
(a)

Discocytes



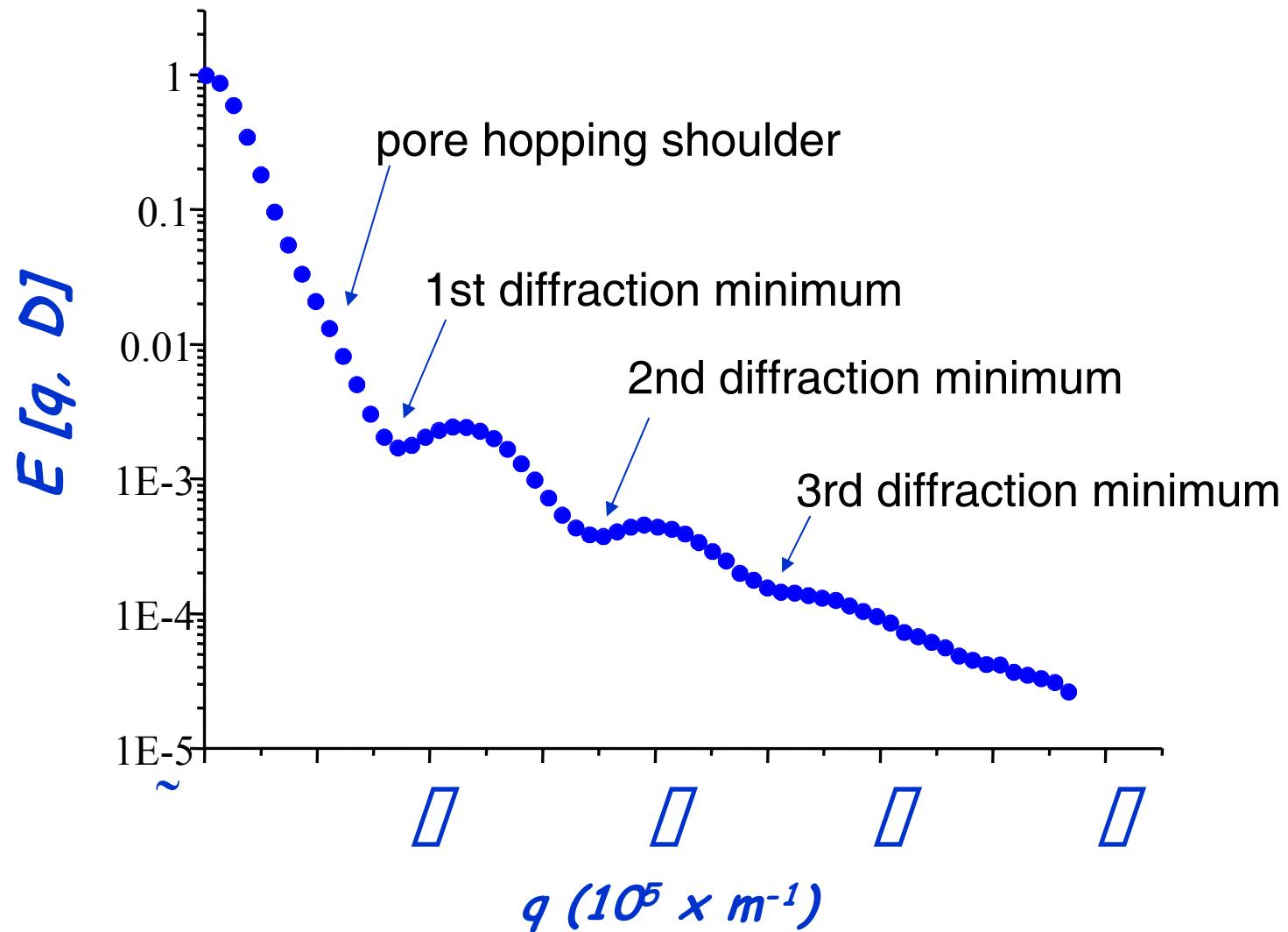
(b)

Spherocytes



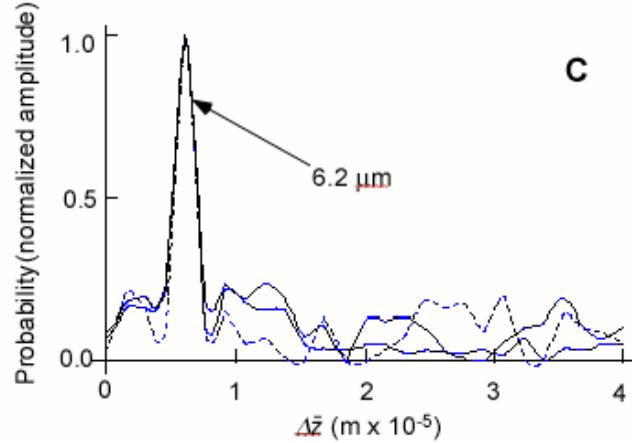
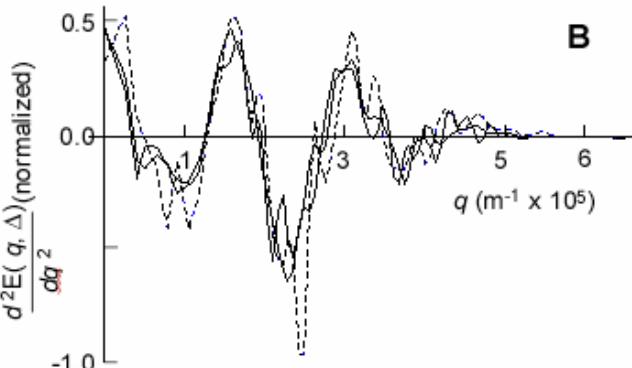
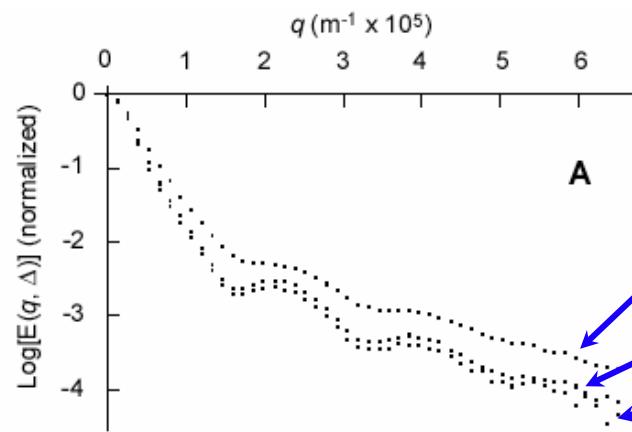


*Tracking red cell shape transformation
in packed samples over time courses of
minutes*



University of Sydney





Ht = 58%
48%
41%

Second
derivative

Fourier
transform

Unbalanced bipolar pulses in STE... rapid signal acquisition



QuickTime™ and a
TIFF (Uncompressed) decompressor
are needed to see this picture.



Movie showing change in shape of the q-space
plot
QuickTime™ and a
decompressor
are needed to see this picture.

q



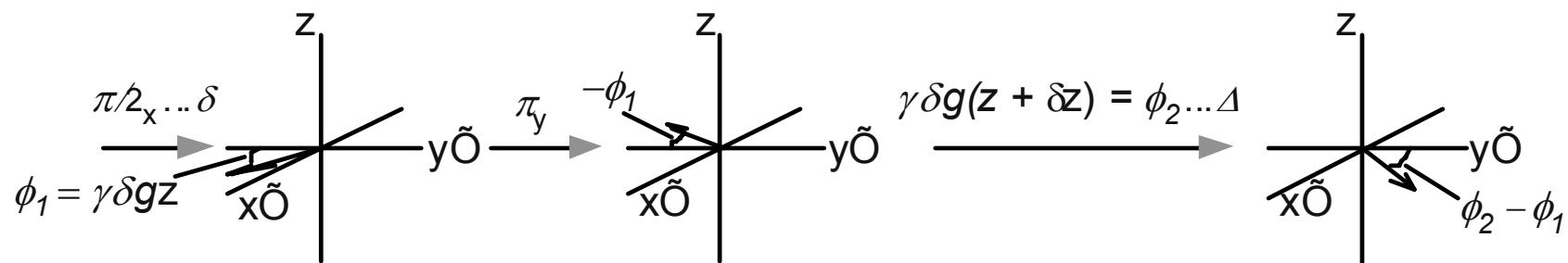
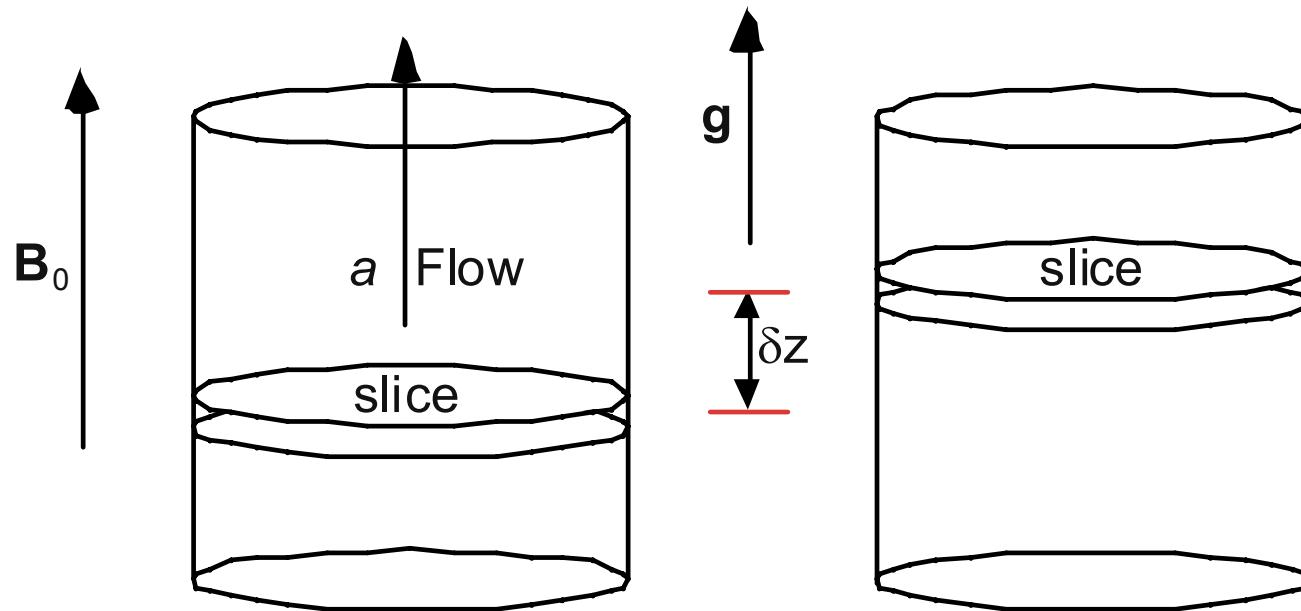
Movie showing change in shape of the mean cell diameter

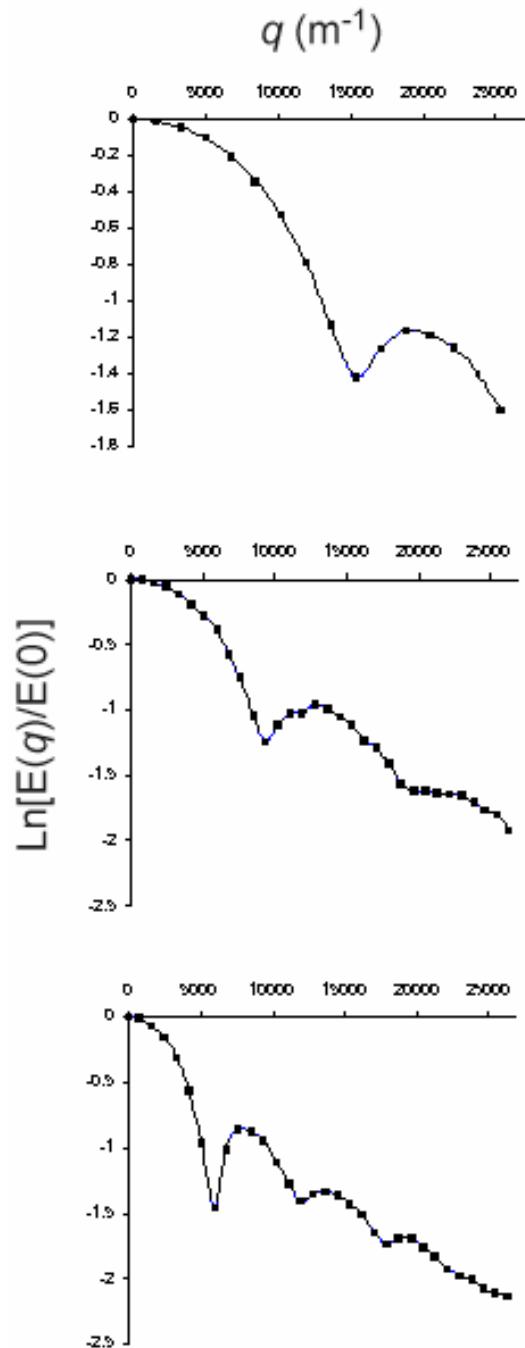
QuickTime™ and a decompressor are needed to see this picture.

z (μm)



Flow diffraction





Flow rate = 0.834 mm s^{-1}

1.32 mm s^{-1}

2.15 mm s^{-1}



Since the uniform linear flow-velocity is $a \text{ m s}^{-1}$ then $a = \delta z / \Delta$; and by definition $\mathbf{q} = (1/2\pi) \gamma \delta \mathbf{g}$ so the change in phase angle brought about by flow is:

$$\begin{aligned}\Delta\phi &= 2\pi q \delta z \\ &= 2\pi q a \Delta\end{aligned}$$



Hence, the normalized signal $S[q, \Delta]$ is proportional to,

$$S[q, \Delta] \propto \sin^2[2\pi q a \Delta] / (2\pi q a \Delta)^2$$

$$q_{\min,n} = n / (2 a \Delta)$$

$$a = n / (2 q_{\min,n} \Delta)$$



$$E[\mathbf{g}, \Delta] = \int \rho[\mathbf{r}] [\mathbf{r} \Delta] \exp[i \delta \mathbf{g} \cdot (\mathbf{r} - \mathbf{r})] d\mathbf{r} d\mathbf{r}$$

$$E[\mathbf{q}, \Delta] = \iint \rho[\mathbf{r}_0] P[\mathbf{r}_0 | \mathbf{r}_0 + \mathbf{R}, \Delta] \exp[i 2 \pi \mathbf{q} \cdot \mathbf{R}] d\mathbf{r}_0 d\mathbf{R}$$

$$\bar{P}[\mathbf{R}, \Delta] = \int \rho[\mathbf{r}_0] P[\mathbf{r}_0 | \mathbf{r}_0 + \mathbf{R}, \Delta] d\mathbf{r}_0$$

$$E[\mathbf{q}, \Delta] = \int \bar{P}[\mathbf{R}, \Delta] \exp[i 2 \pi \mathbf{q} \cdot \mathbf{R}] d\mathbf{R}$$

$$\bar{P}[\mathbf{R}, \Delta] = \int E[\mathbf{q}, \Delta] \exp[-i 2 \pi \mathbf{q} \cdot \mathbf{R}] d\mathbf{q}$$

$$E[\mathbf{q}, \infty] = \int \rho[\mathbf{r}_0] \exp[i 2 \pi \mathbf{q} \cdot \mathbf{r}_0] d\mathbf{r}_0 \int \rho[\mathbf{r}] \exp[i 2 \pi \mathbf{q} \cdot \mathbf{r}] d\mathbf{r}$$

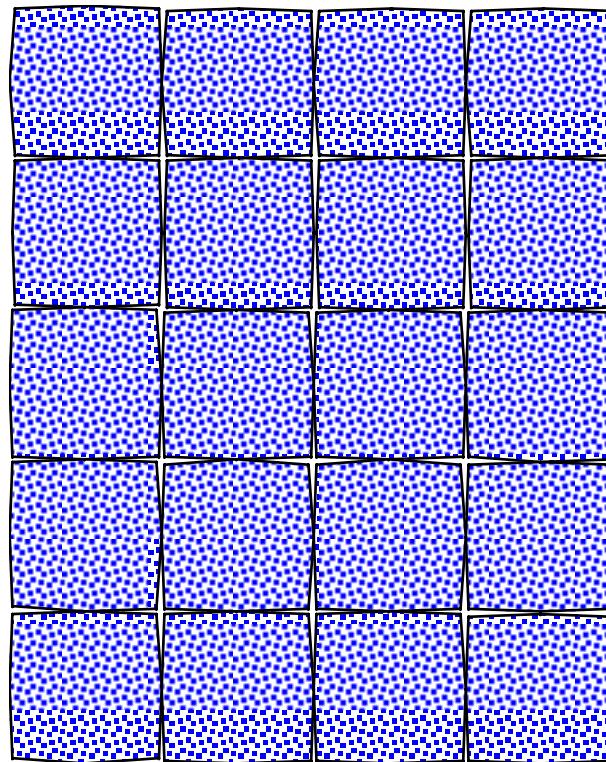
$$E[\mathbf{q}, \infty] = S^*[\mathbf{q}] S[\mathbf{q}] = |S[\mathbf{q}]|^2$$

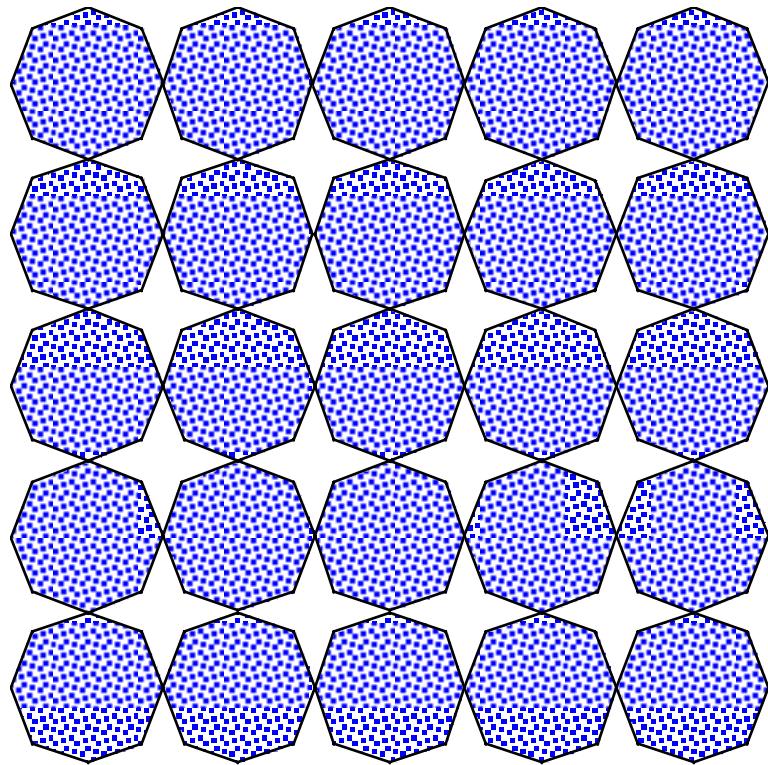


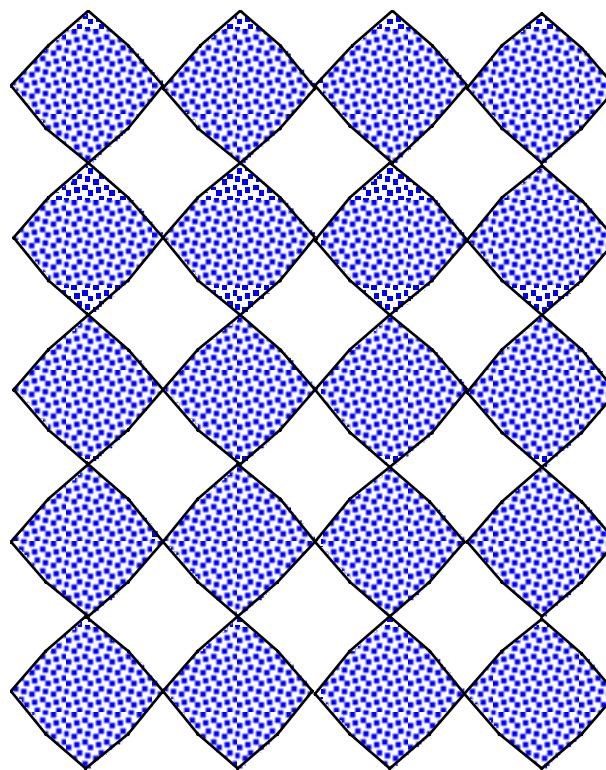
A model of variable packing density

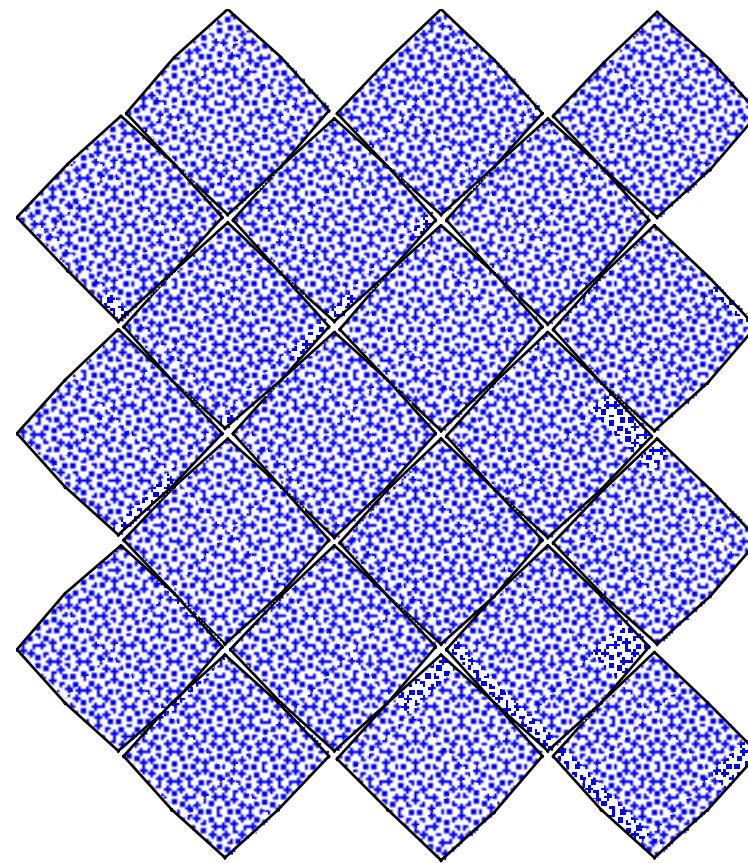
q-space plot results

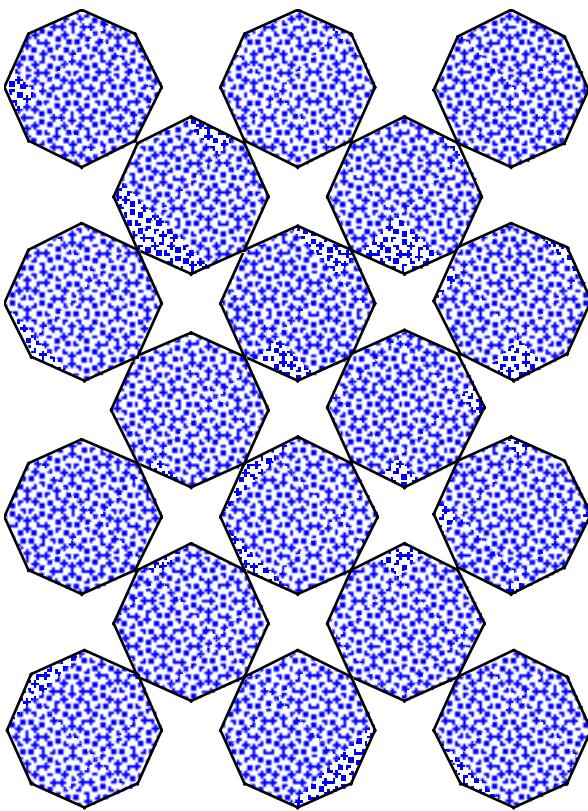
...octagon-star system

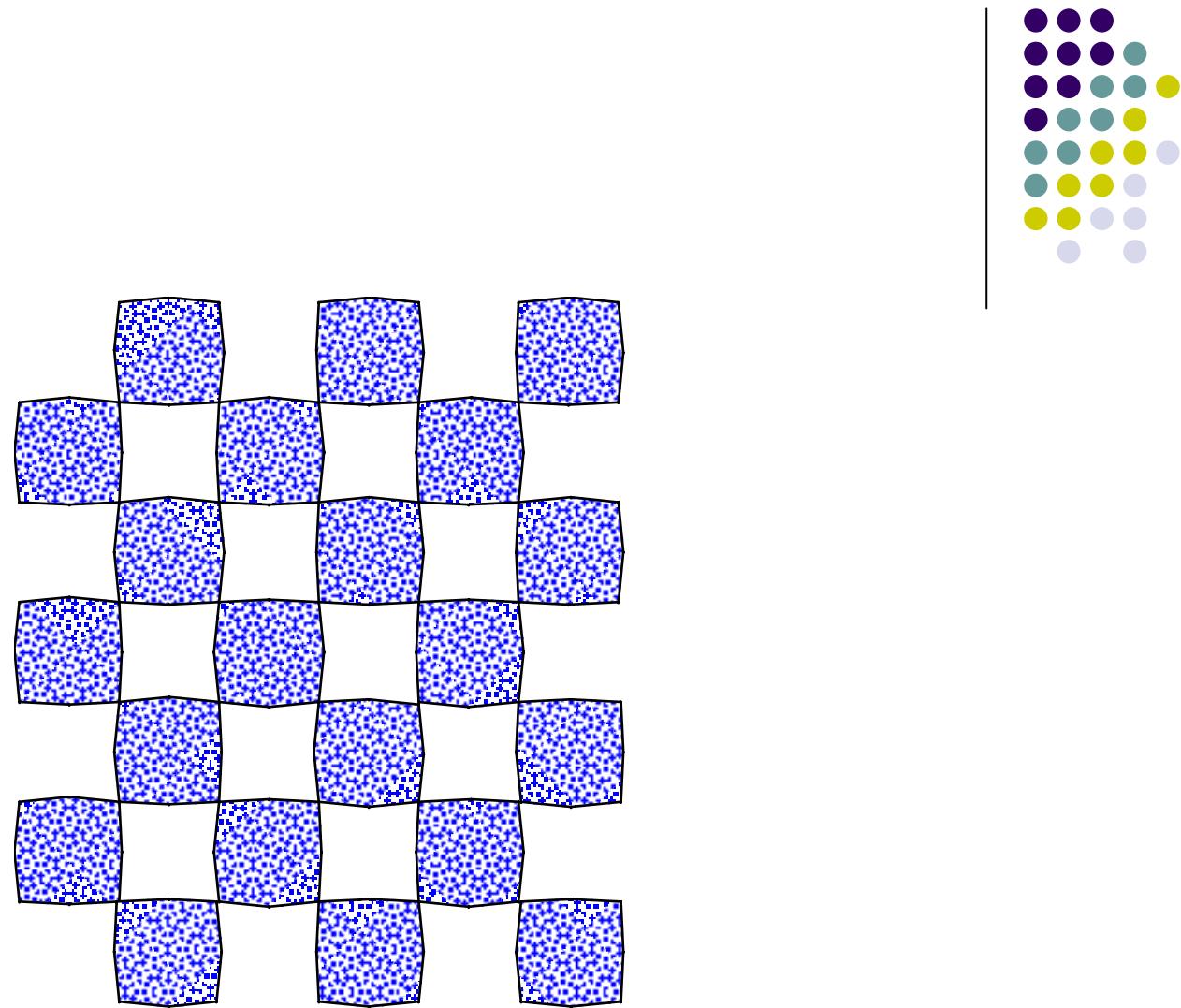






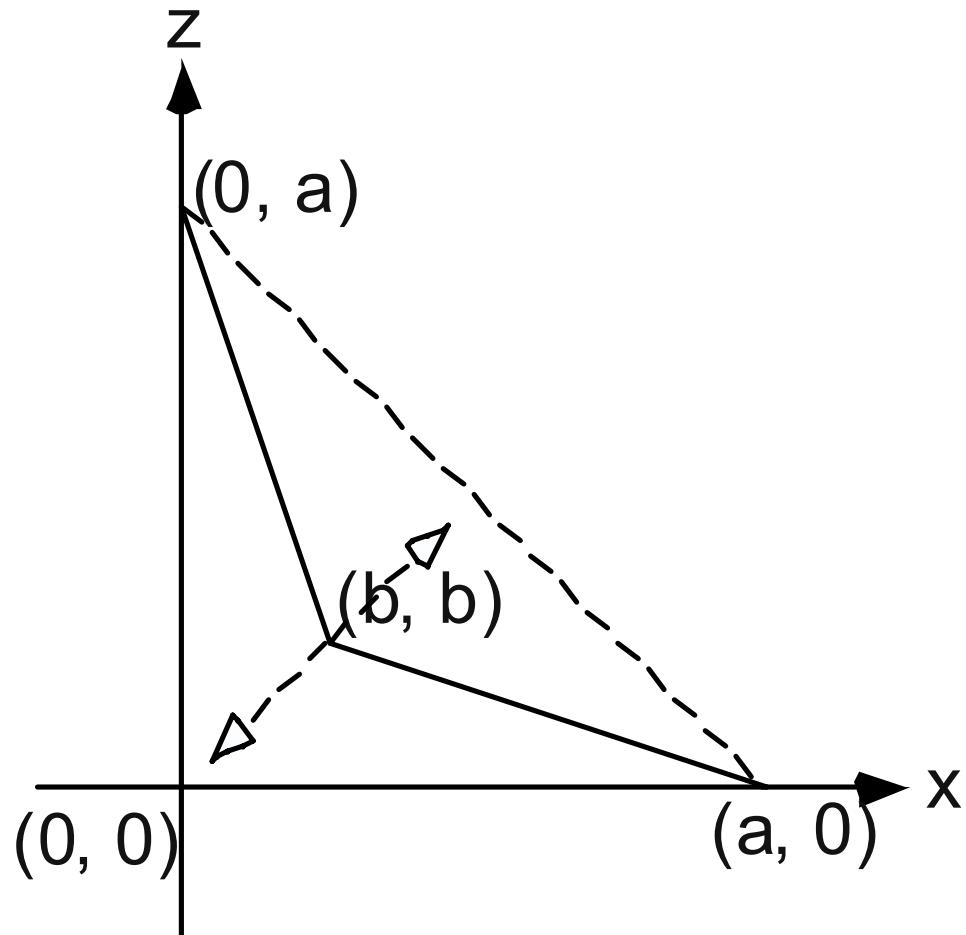




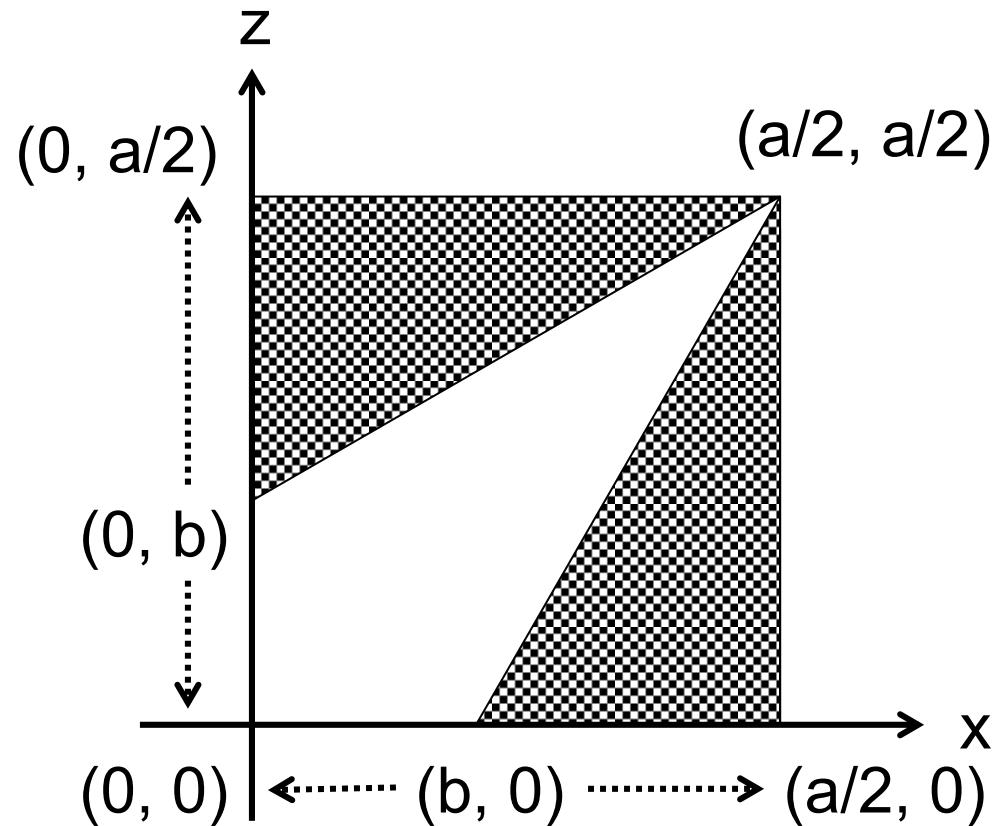


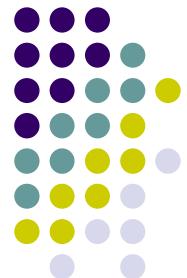


Construction element for q -space signal intensity for octagon-star model



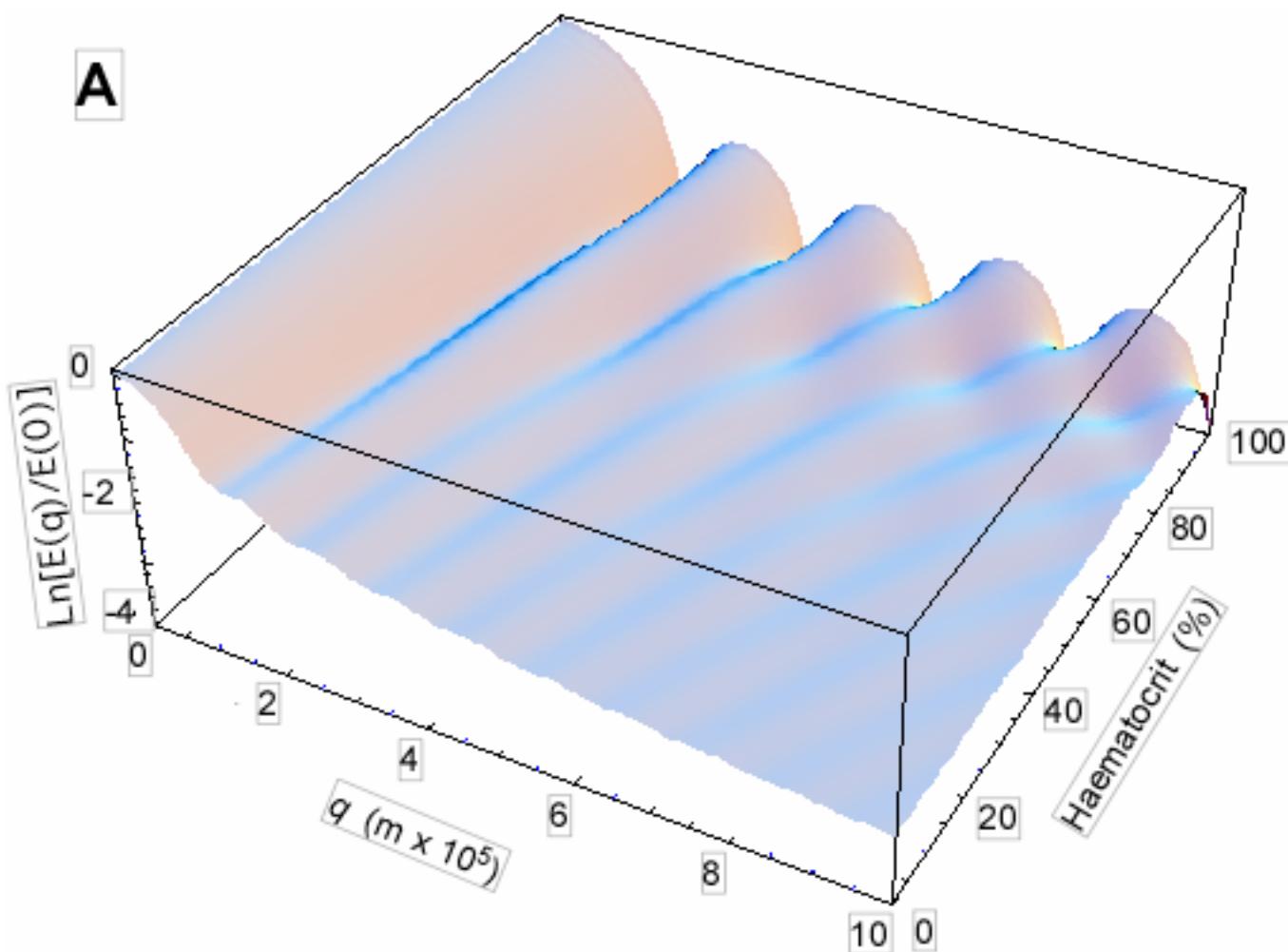
Construction element for q -space signal intensity
for octagon-star model...90° rotation

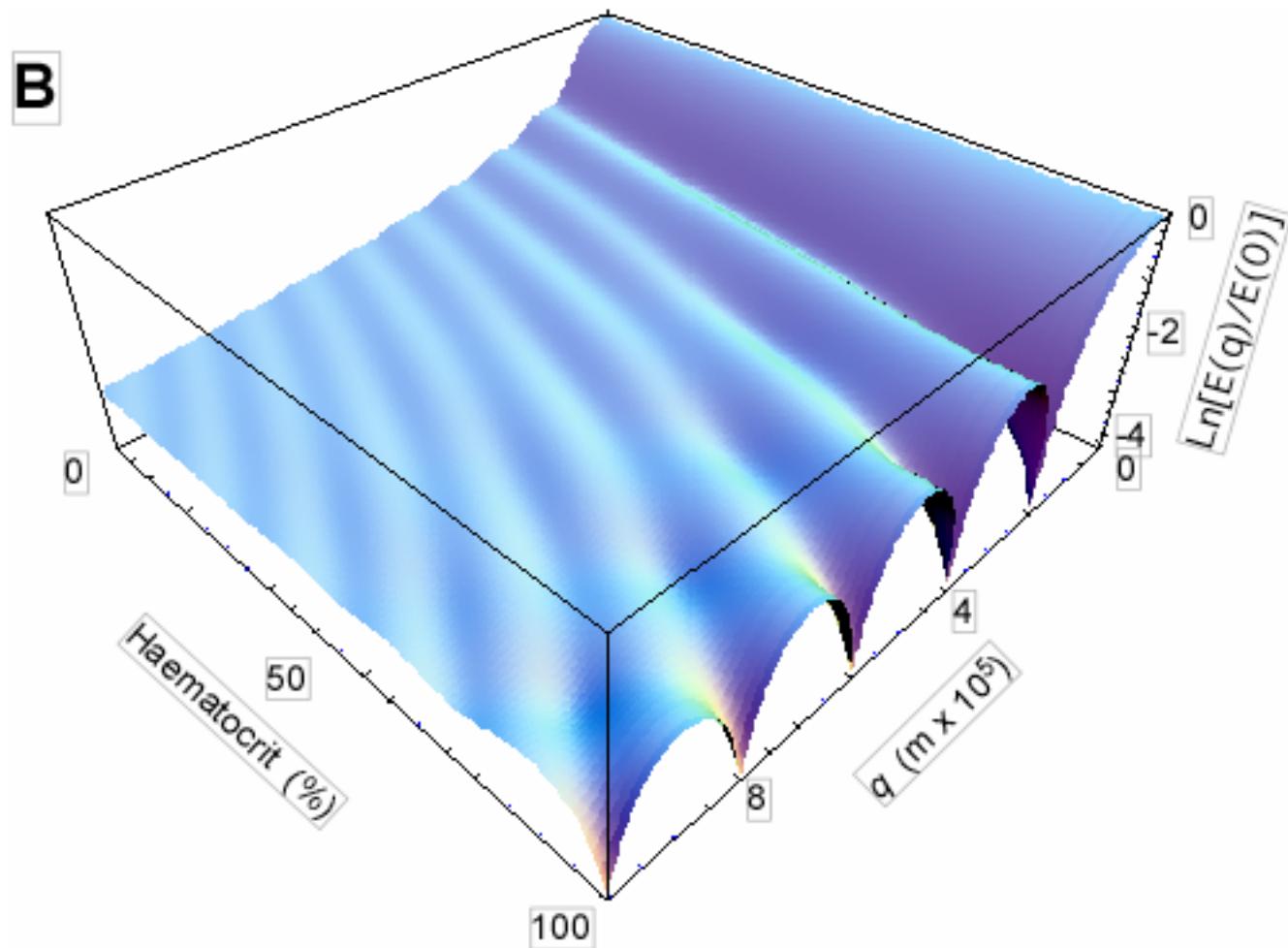


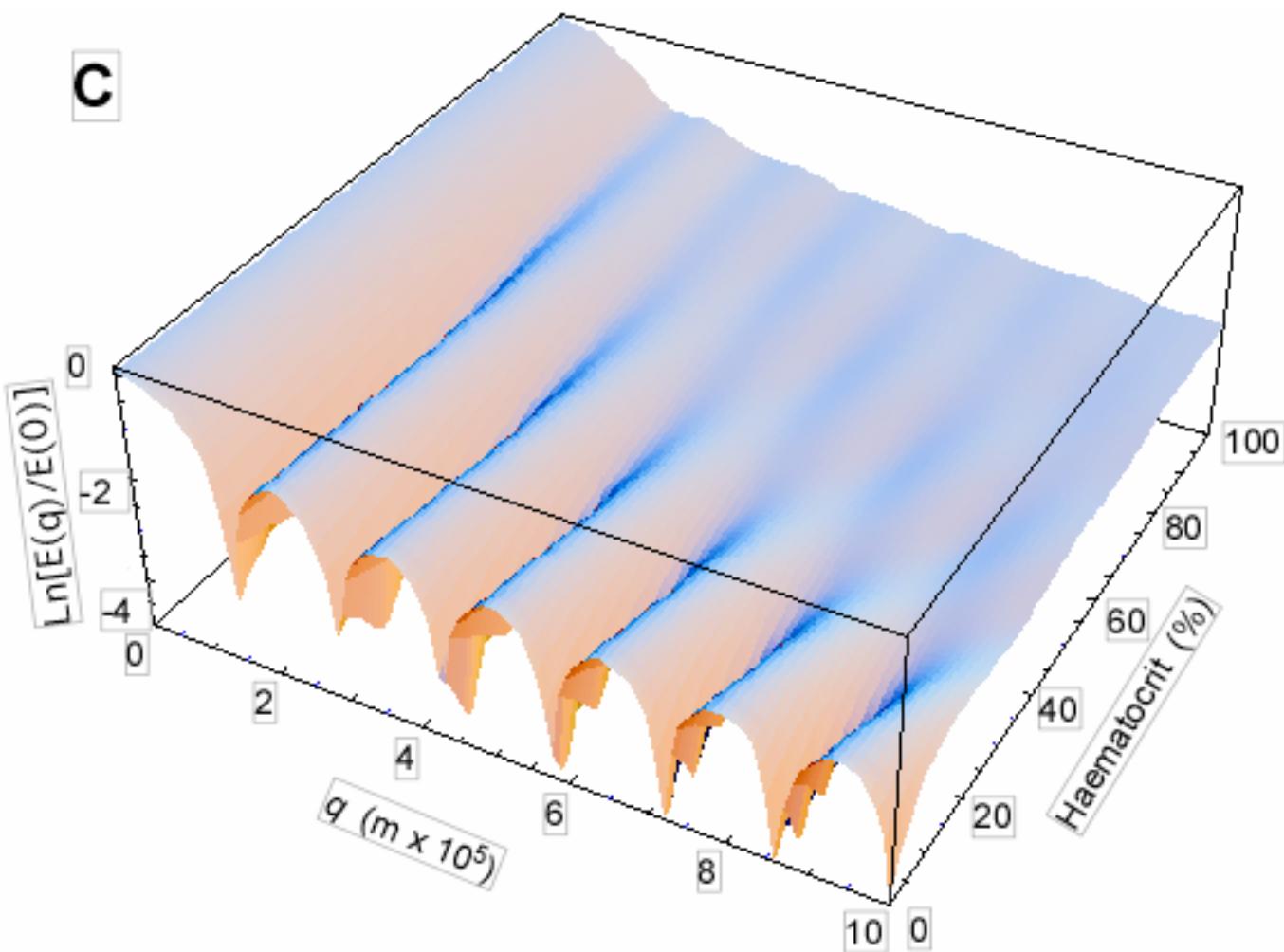


Expression for q -space signal intensity from stars alone

$$E_{star}[q, \infty] = \frac{1}{8a^2(a-2b)^2\pi^4 q^4} \times$$
$$\left(a^4 - 4a^3b + 20a^2b^2 - 32ab^3 + 16b^4 + 2a^2(a-2b)^2b^2\pi^2q^2 - a^2(a-2b)^2\cos[2b\pi q] + \right.$$
$$4(a-b)b\left((a-2b)^2\cos[a\pi q] - a(\cos[(a-2b)\pi q] + 2(a-2b)b\pi q \sin[a\pi q]) \right) +$$
$$\left. 2a^3(a-2b)b\pi q \sin[2b\pi q] \right)$$





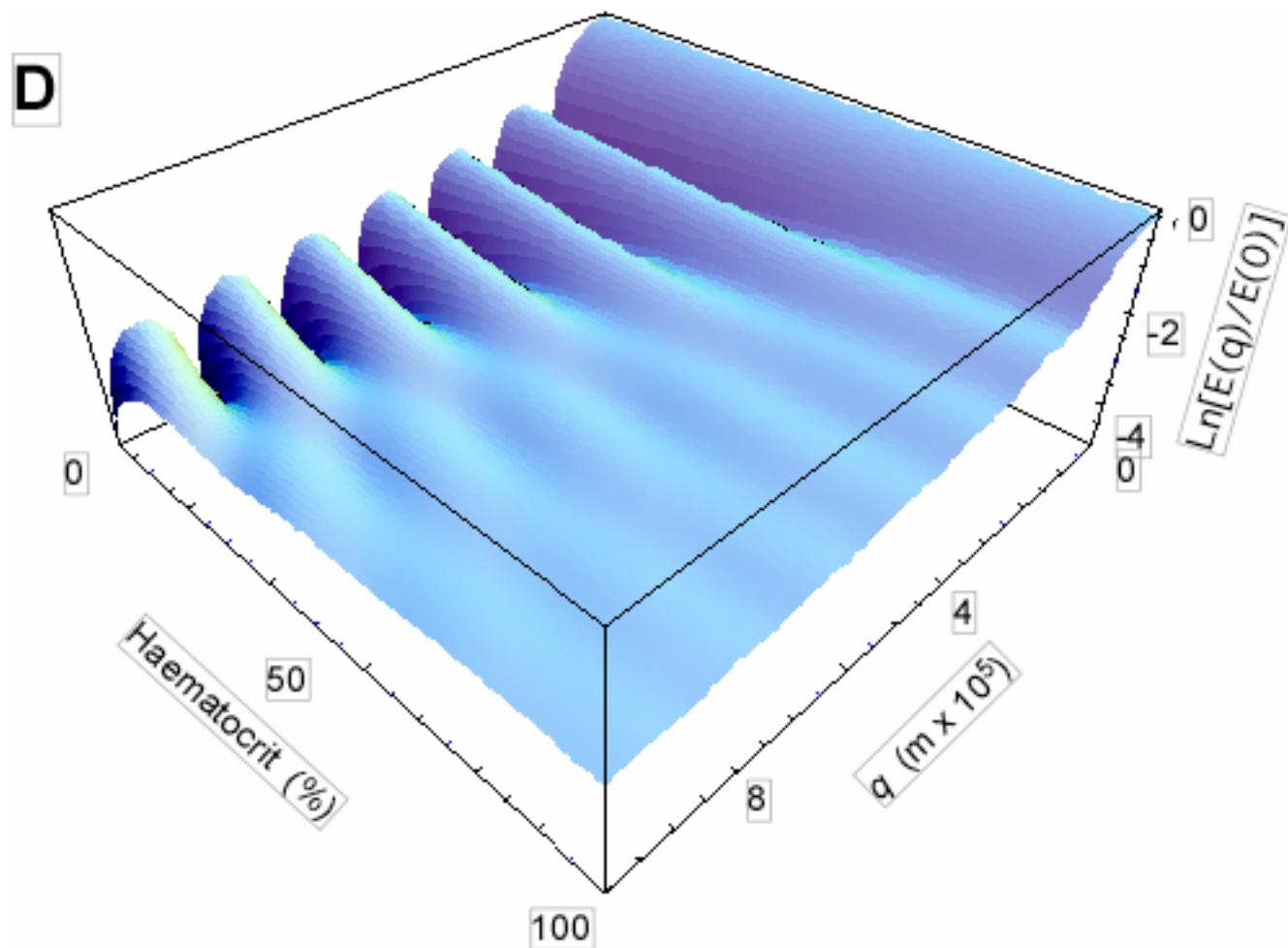
**C**

q -space plot from octagon-star system as H_t increases

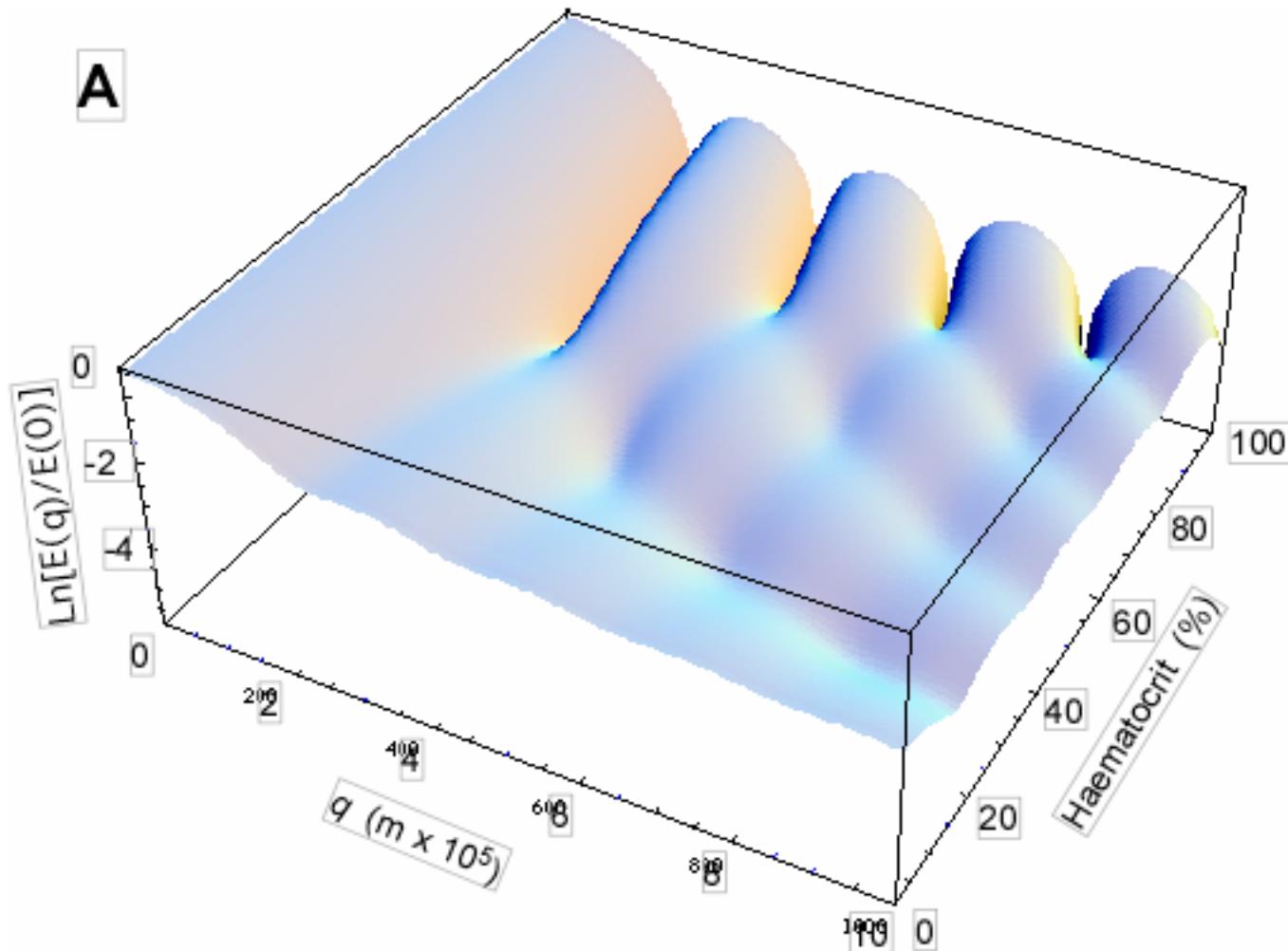
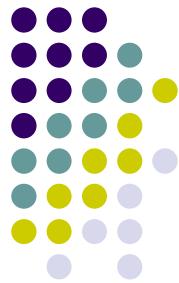


QuickTime™ and a decompressor are needed to see this picture.

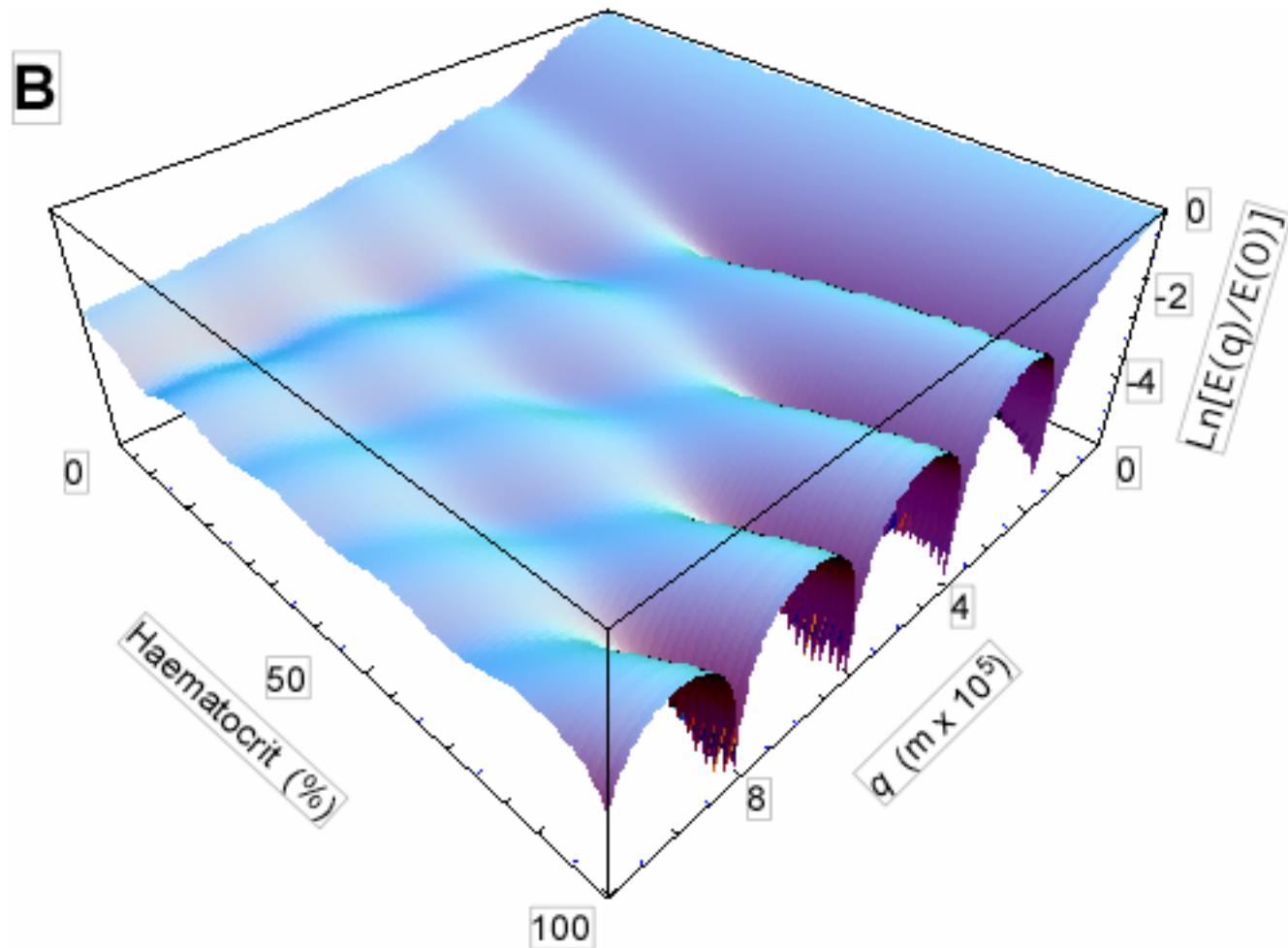
Movie showing change in shape of the q -space plot-function as H_t is increased



q -space plot from stars only



q-space plot from stars only



Thanks to past and present students...



TIFF (Uncompressed) files
are needed to see this picture.

TIFF (Uncompressed) files
are needed to see this picture.

QuickTime™ and a
TIFF (Uncompressed) decompressor
are needed to see this picture.

Bill Price is my name!





Bill Bubb	NMR
Bob Chapman	NMR
Tom Eykyn	NMR
David Jacques	DIC
Tim Larkin	<i>Mathematica</i>
Guilhem Pages	NMR & fast <i>q</i> -space
David Regan	NMR & simulation
David Szekely	DIC & <i>Mathematica</i>

Chris Garvey
Bill Price
Peter Stilbs

Grazie!

School of Molecular and Microbial Biosciences University of Sydney

